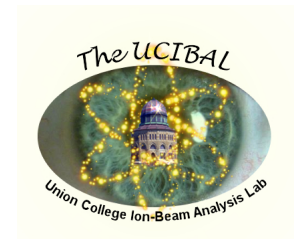
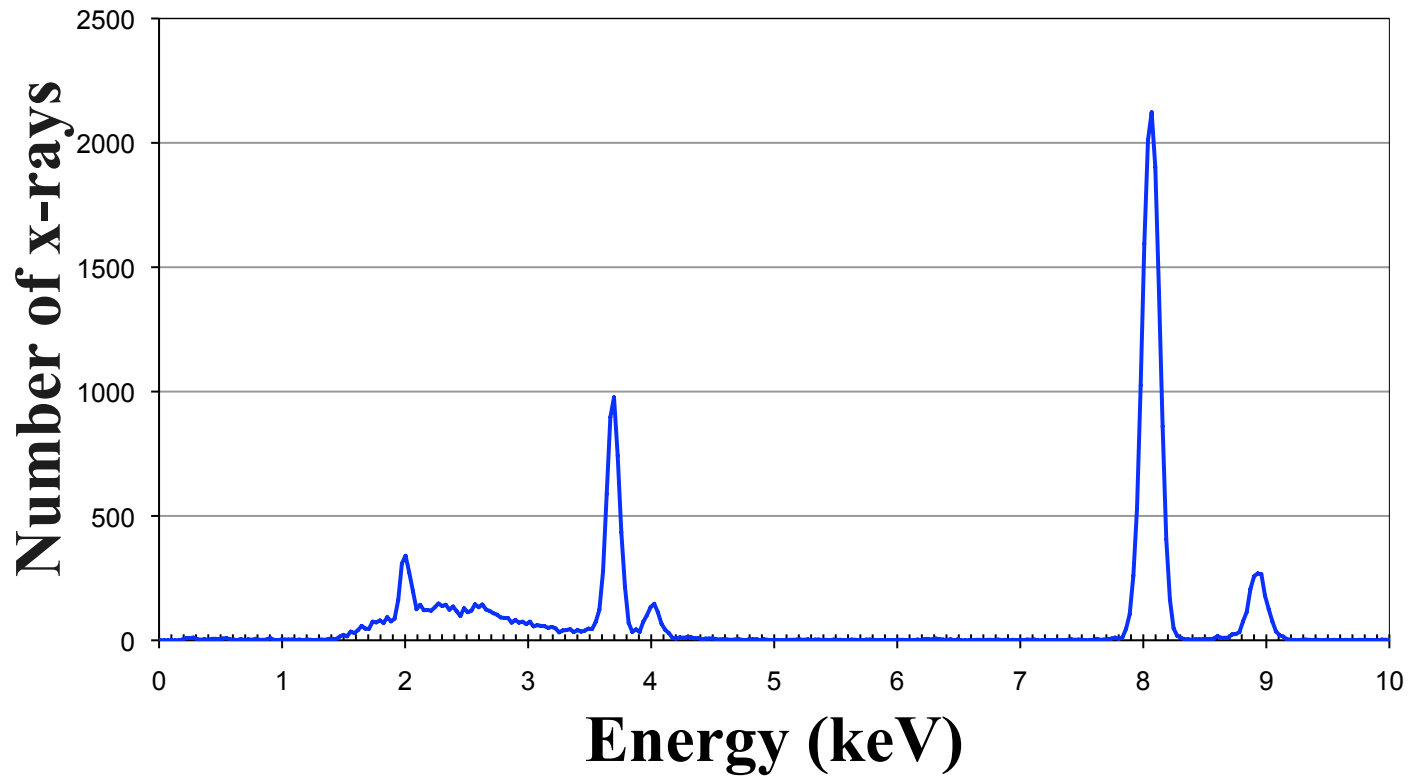


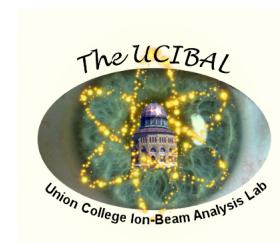
Calculations of Characteristic X-ray Energies and Wavelengths and the PIXE Spectrum



The Characteristic X-ray Wavelengths

- **Electronic transitions within inner shells of heavier atoms are accompanied by large energy transfers.**
- **The inner electrons of high Z elements are bound tightly to the atom, since they see essentially the entire nuclear charge.**
- **First let's make do a small calculation in order to simplify or lives when we calculate the energies of the orbits.**

$$\begin{aligned} E_n &= -\frac{Z^2 m e^4}{2(4\pi\epsilon_0)^2 n^2 \hbar^2} = -\left(\frac{m e^4}{2(4\pi\epsilon_0)^2 \hbar^2}\right) \frac{Z^2}{n^2} \\ &= -\left[\frac{(9.11 \times 10^{-31} \text{ kg})(1.6 \times 10^{-19} \text{ C})^4}{32\pi^2 (8.85 \times 10^{-12} \frac{\text{C}^2}{\text{Nm}^2}) \left(\frac{6.63 \times 10^{-34} \text{ Js}}{2\pi}\right)^2} \times \frac{1 \text{ eV}}{1.6 \times 10^{-19} \text{ J}} \right] \frac{Z^2}{n^2} \\ &= -(13.57 \text{ eV}) \frac{Z^2}{n^2} \end{aligned}$$



The Characteristic X-ray Wavelengths

- Let's calculate the lowest energy, or longest expected x-ray wavelengths for the element copper.

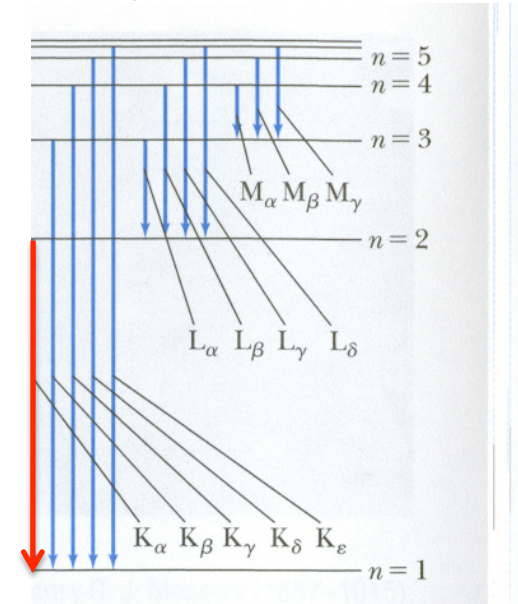


- The energy of an inner shell electron is given by $Z = 29$, and $n = 1$.

$$E_1 = -(13.57\text{eV}) \frac{Z^2}{n_{\text{lower}}^2} = -(13.57\text{eV}) \frac{(29)^2}{(1)^2} = -11412.4\text{eV}$$

- The energy of an outer shell electron is given by $Z = 29$, and $n = 2$.

$$E_2 = -(13.57\text{eV}) \frac{Z^2}{n_{\text{upper}}^2} = -(13.57\text{eV}) \frac{(29)^2}{(2)^2} = -2853.1\text{eV}$$



Thornton, S., & Rex, A., Modern Physics for Scientists and Engineers, 3rd Ed., Thomas Brooks Cole 151(2006).

- This is the transition of an electron from the $n = 2$ state to the $n = 1$ state, or an electronic transition from the L-shell to the K-shell.

- This transition is called the K_{α} transition for copper and the difference in energy between these states is the energy of the emitted x-ray.



The Characteristic X-ray Wavelengths

The energy of the emitted photon is the difference in energy between the upper state ($n = 2$) and the lower state ($n = 1$).

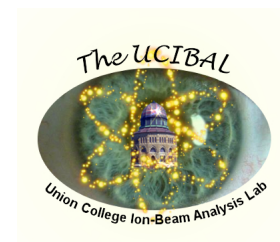
$$\Delta E = E_{upper} - E_{lower} = -2853.1eV - (-11412.4eV) = 8559.3eV$$

This corresponds to a wavelength of

$$\Delta E = \frac{hc}{\Delta\lambda}$$
$$\Delta\lambda = \frac{hc}{\Delta E} = \frac{\left(6.63 \times 10^{-34} \text{ Js} \times \frac{1eV}{1.6 \times 10^{-19} \text{ J}}\right) 3 \times 10^8 \frac{m}{s}}{8559.3eV} = 1.45 \times 10^{-10} \text{ m}$$

The actual wavelength (measured in the laboratory) is $1.54 \times 10^{-10} \text{ m}$.

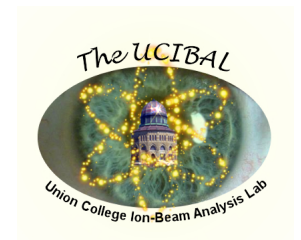
This is about 70% from the true value!! Hmm...



The Characteristic X-ray Wavelengths

Comments:

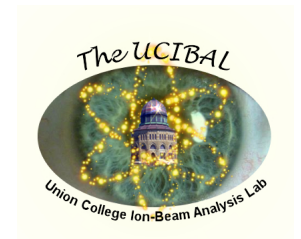
- These wavelengths are calculated based on a hydrogen-like atom using the Bohr model of the atom.
- This means that there is a single electron that transitions.
- The problem with heavy or high Z atoms is that they are rarely single electron atoms.
- So do we live with this or can we fix our results and theory?
- I guess we have to fix the theory since it doesn't give the expected results.



Modifications to the Bohr Theory

- To start, in multi-electron atoms the higher orbital electrons are partially screened from the nucleus.
- In other words they don't see the full nuclear charge of the nucleus.
- The net charge an electron say in the L -shell sees is Ze^- due to the nucleus *minus* e^- due to the one electron in the K -shell (one was ejected.)
- Now, this is only the L to K -shell transitions. Further modifications are needed from M to L -shell transitions, for example.
- Therefore the net charge is $(Z - 1)e^-$

- The potential energy is thus
$$V_n = -\frac{(Z - 1)e^2}{4\pi\epsilon_0 r_n}$$



Modifications to the Bohr Theory

- The energy of the orbit is given as

$$E'_n = \frac{1}{2}mv_n^2 + V_n = \frac{1}{2}m\left(\frac{(Z-1)e^2}{4\pi\epsilon_0 n\hbar}\right)^2 - \frac{(Z-1)e^2}{4\pi\epsilon_0 \frac{4\pi\epsilon_0 n^2 \hbar^2}{m(Z-1)e^2}}$$

- Doing the math...

$$E_n = -\frac{(Z-1)^2 me^4}{2(4\pi\epsilon_0)^2 n^2 \hbar^2}$$

- This is the modified Bohr theory to take into account screening of the outer shell electrons by the inner shell electrons.

- How do our calculations look now? Did we do any better?



Modifications to the Bohr Theory and the new X-ray Wavelengths

- Let's recalculate the expected x-ray energy and wavelength for the K_α transition in copper.

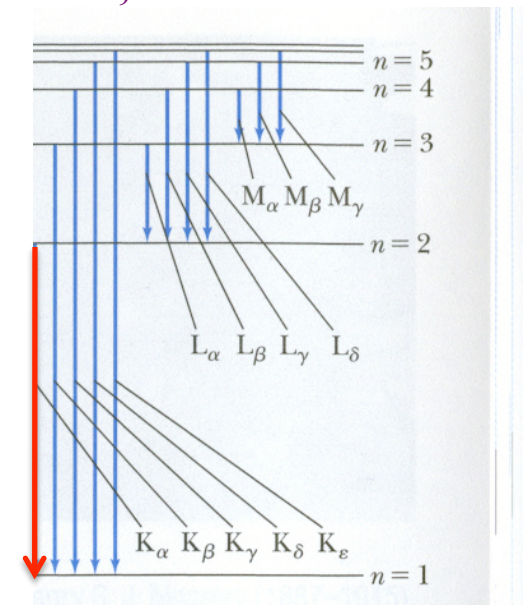


- The energy of an inner shell electron is given by $Z = 29$, and $n = 1$.

$$E_1 = -(13.57\text{eV}) \frac{(Z-1)^2}{n_{\text{lower}}^2} = -(13.57\text{eV}) \frac{(28)^2}{(1)^2} = -10638.9\text{eV}$$

- The energy of an outer shell electron is given by $Z = 29$, and $n = 2$.

$$E_2 = -(13.57\text{eV}) \frac{(Z-1)^2}{n_{\text{upper}}^2} = -(13.57\text{eV}) \frac{(28)^2}{(2)^2} = -2659.7\text{eV}$$



- This is the transition of an electron from the $n = 2$ state to the $n = 1$ state, or an electronic transition from the L-shell to the K-shell including screening.

Modifications to the Bohr Theory and the new X-ray Wavelengths

The energy of the emitted photon is the difference in energy between the upper state ($n = 2$) and the lower state ($n = 1$).

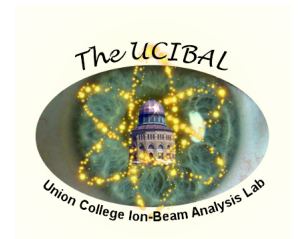
$$\Delta E = E_{upper} - E_{lower} = -2659.7eV - (-10638.9eV) = 7979.2eV$$

This corresponds to a wavelength of

$$\Delta E = \frac{hc}{\Delta\lambda}$$
$$\Delta\lambda = \frac{hc}{\Delta E} = \frac{\left(6.63 \times 10^{-34} \text{ Js} \times \frac{1eV}{1.6 \times 10^{-19} \text{ J}}\right) 3 \times 10^8 \frac{m}{s}}{7979.2eV} = 1.56 \times 10^{-10} \text{ m}$$

The actual wavelength (measured in the laboratory) is $1.54 \times 10^{-10} \text{ m}$.

This is about 1.2% from the true value!!



More X-ray Wavelengths...

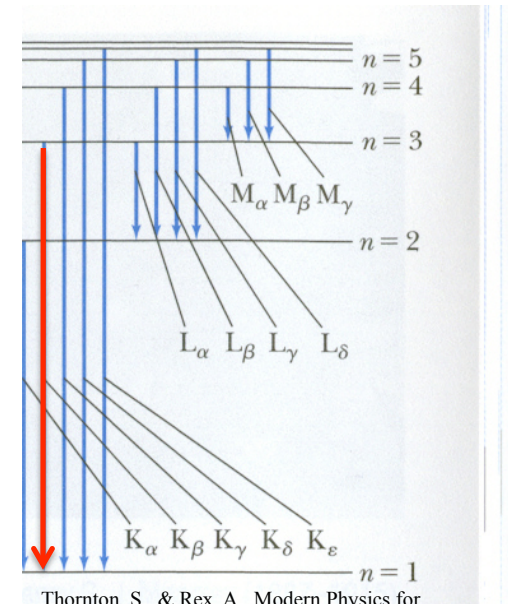
- What is the K_β wavelength for Copper?



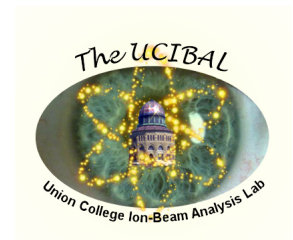
- Recalling the energy level diagram for an atom, the K_β transition is from the $n = 3$ state to the $n = 1$ state.
- The energies of the upper and lower states are thus

$$E_1 = -(13.57\text{eV}) \frac{(Z-1)^2}{n_{\text{lower}}^2} = -(13.57\text{eV}) \frac{(28)^2}{(1)^2} = -10638.9\text{eV}$$

$$E_2 = -(13.57\text{eV}) \frac{(Z-1)^2}{n_{\text{upper}}^2} = -(13.57\text{eV}) \frac{(28)^2}{(3)^2} = -1182.1\text{eV}$$



Thornton, S., & Rex, A., Modern Physics for Scientists and Engineers, 3rd Ed., Thomas Brooks Cole 151(2006).



More X-ray Wavelengths...

The energy of the emitted photon is the difference in energy between the upper state ($n = 3$) and the lower state ($n = 1$).

$$\Delta E = E_{upper} - E_{lower} = -1182.1eV - (-10638.9eV) = 9456.8eV$$

This corresponds to a wavelength of

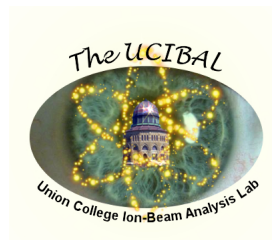
$$\Delta E = \frac{hc}{\Delta\lambda}$$
$$\Delta\lambda = \frac{hc}{\Delta E} = \frac{\left(6.63 \times 10^{-34} \text{ Js} \times \frac{1eV}{1.6 \times 10^{-19} \text{ J}}\right) 3 \times 10^8 \frac{m}{s}}{9456.8eV} = 1.32 \times 10^{-10} \text{ m}$$

The actual wavelength (measured in the laboratory) is $1.39 \times 10^{-10} \text{ m}$.

This is about 5.3% from the true value!!

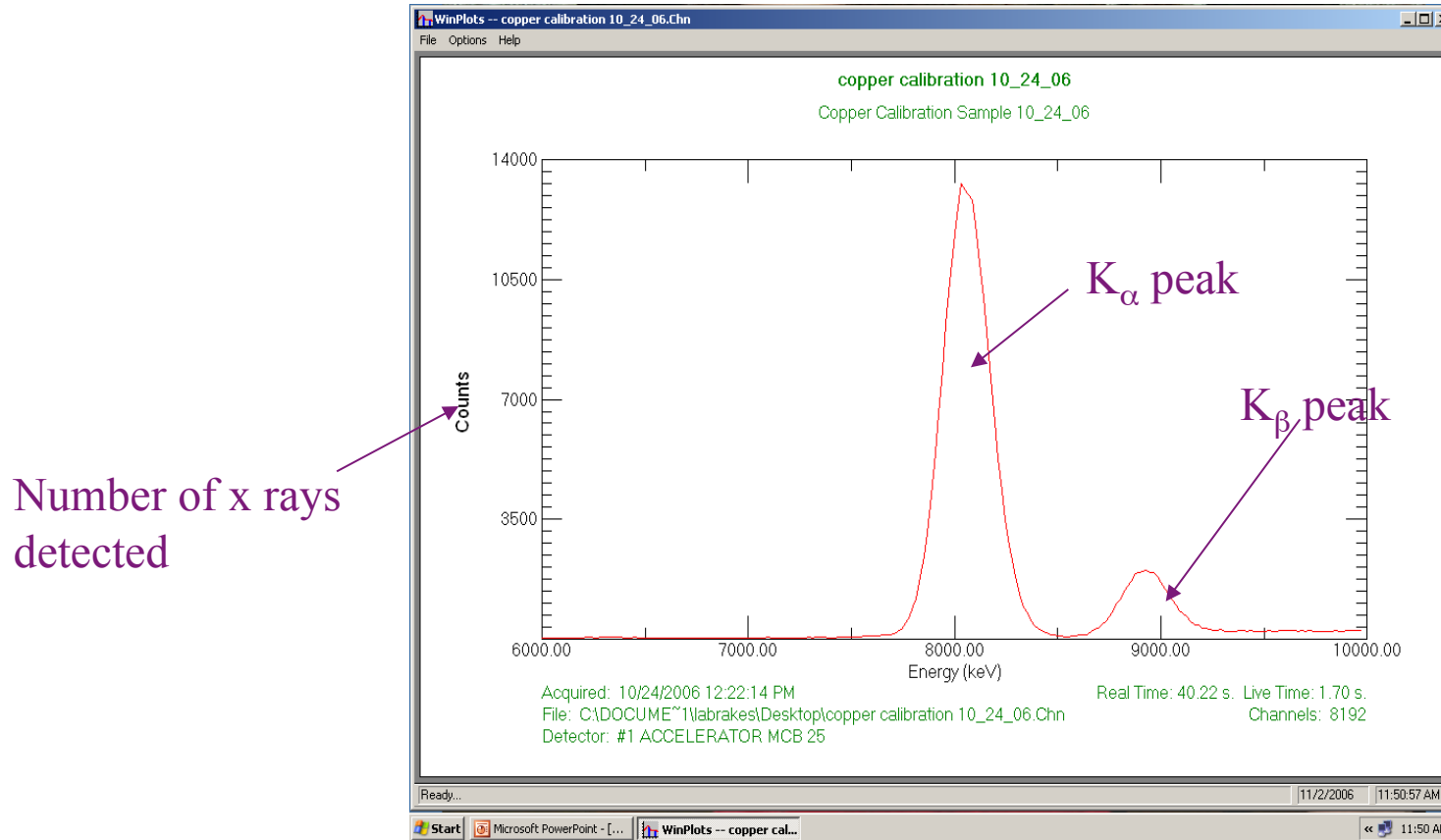


These energies are actually tabulated and in practice look up the element based on the transition energies observed on a calibrated x-ray spectrum

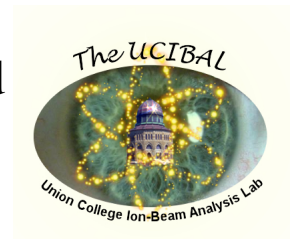


The X-ray Spectrum of Copper

This is a typical *PIXE* plot that shows the x-ray spectrum of copper.

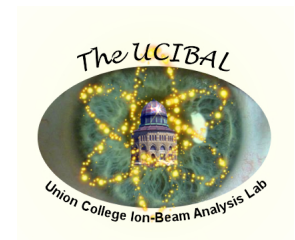
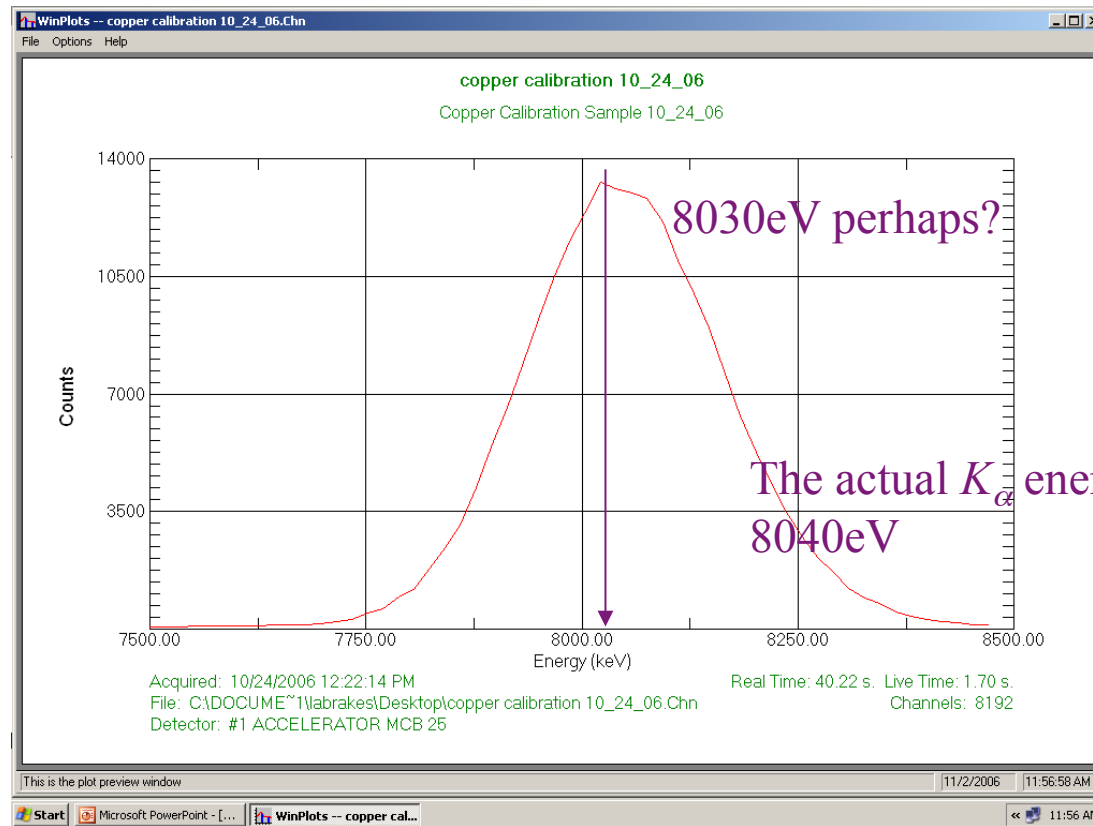


This sample was run on our accelerator and calibrated using a set of standards.

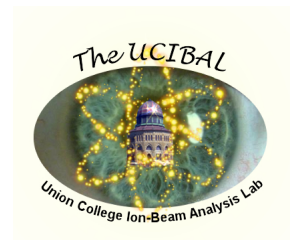
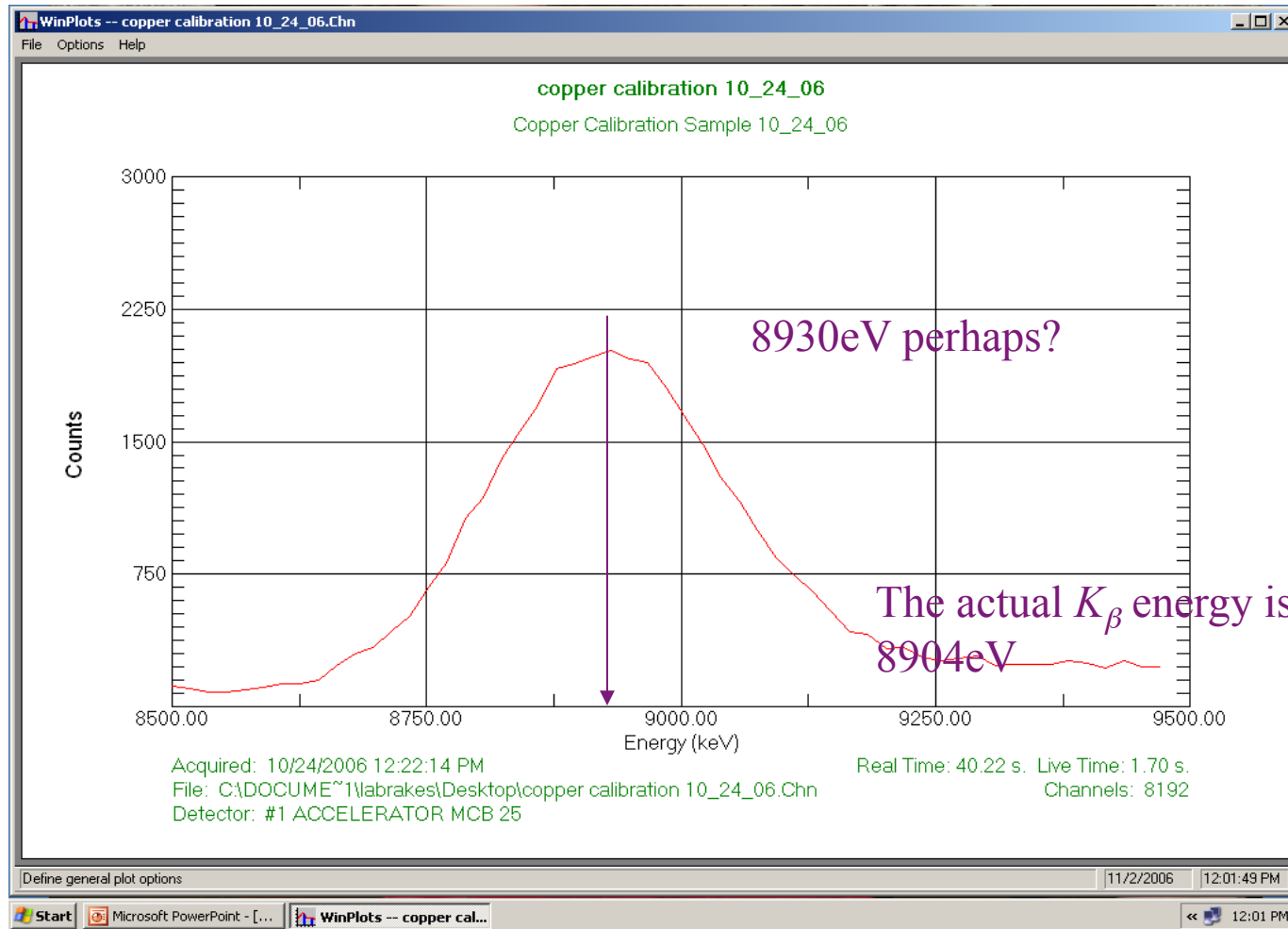


The X-ray Spectrum of Copper

Here we will read off the peak energies and compare those experimentally determined peak energies with the energies of the transitions that we just calculated.



The X-ray Spectrum of Copper



Conclusions

- So, we can calculate the x-ray transition energies to a fairly high degree of accuracy.
- There are lots of other effects we haven't looked at, absorption of x-ray, attenuation of x-rays, failure to produce an x-ray (Auger electrons)...
- Screened Bohr model seems to work well to describe the transitions.
- X-ray energies for K-series transitions scale with $(Z - 1)^2$.
- L-series x-rays are more complicated how do we describe them?
- Further, how much of the elements are present?
- What are the environmental sources of the elements you found?
- What is the chemical identity of the elements – what are the elements bonded too?

